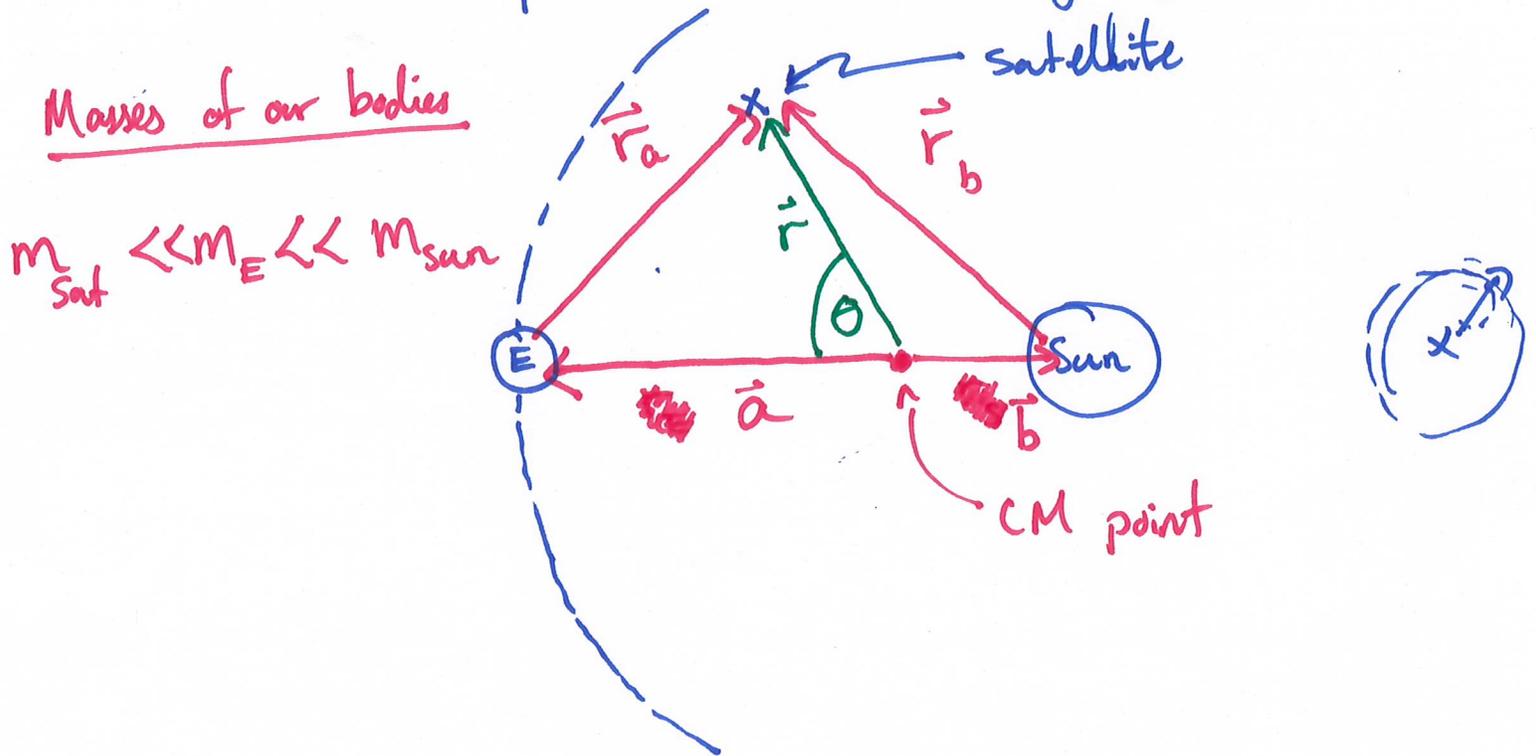


Lagrange Points : Part 1

- stable points in a 3-body system:



A stable point means that

- ① zero net force at that point
- ② configuration (when viewed in a rotating reference frame) does not change.

↳ Around what point does the system rotate?

Strategy: ① Calculate effective potential
 $V(r, \theta)$

includes centrifugal force
(in rotating frame)

② take derivatives to
solve for \vec{F}

③ set $\vec{F} = 0$
defines stability.

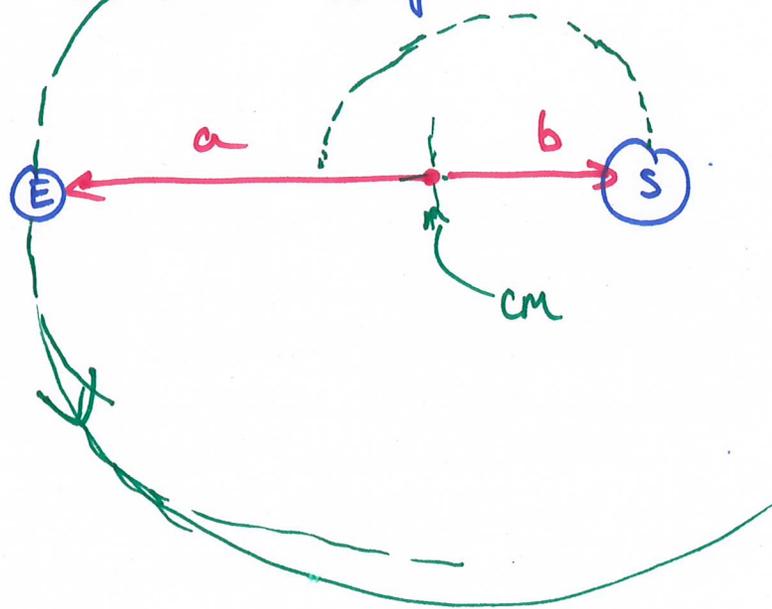
① Effective potential:

$$V(r, \theta) = -\frac{GM_s m}{r_b} - \frac{GM_E m}{r_a} + \frac{1}{2} m r^2 \omega_0^2$$


What is ω for this system?

$$\omega_0 = \omega_{\text{Earth-Sun system}}$$

To calculate ω_0 , we can use a similar setup.



Relationship between a & b :

$$m_E a = m_S b$$

$$b = \frac{M_E}{M_S} a = \alpha a$$

$$\alpha = \frac{M_E}{M_S} \ll 1 \quad (\text{small \#})$$

Force analysis for the Earth:

radial force balance

$$F_r = \frac{-G M_S M_E}{(a+b)^2} + M_E \omega_0^2 a = 0$$

$$\omega_0^2 = \frac{GM_s}{a(a+b)^2}$$

$$b = \alpha a$$

$$\omega_0^2 = \frac{GM_s}{a^3(1+\alpha)^2}$$

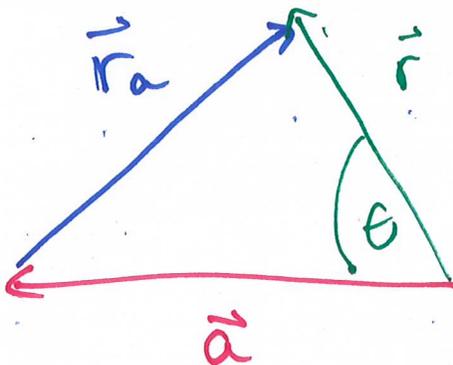
To calculate stability, we need the forces. \Rightarrow take derivatives of the potential.

$$\frac{\partial V}{\partial \theta} = 0$$

$$\text{and } \frac{\partial V}{\partial r} = 0$$

Need $r_a = r_a(r, \theta)$

$$r_b = r_b(r, \theta)$$



$$\vec{r} = \vec{a} + \vec{r}_a$$

$$\vec{r}_a = \vec{r} - \vec{a}$$

Law of cosines:

$$\vec{r}_a \cdot \vec{r}_a = r_a^2$$

$$(\vec{r} - \vec{a}) \cdot (\vec{r} - \vec{a}) = \vec{r} \cdot \vec{r} - \vec{r} \cdot \vec{a} \\ - \vec{a} \cdot \vec{r} + \vec{a} \cdot \vec{a}$$

$$\vec{a} \cdot \vec{r} = \vec{r} \cdot \vec{a}$$

$$\vec{a} \times \vec{r} = -\vec{r} \times \vec{a}$$

$$= r^2 - 2\vec{r} \cdot \vec{a} + a^2$$

$$= r^2 + a^2 - 2ra \cos(\theta)$$

$$r_a = \left(r^2 + a^2 - 2ra \cos(\theta) \right)^{1/2}$$

$$r_b = \left(r^2 + b^2 + 2rb \cos(\theta) \right)^{1/2}$$

$$= \left(r^2 + b^2 - 2rb \underbrace{\cos(180-\theta)}_{-\cos(\theta)} \right)^{1/2}$$