

Radiation for Heat transfer rate

Consider a lump of charcoal in space at a temperature T . The power required to keep this lump of matter @ this temperature is the radiative heat transfer rate, \dot{q}_r . This

depends on 3 things

- the surface area of the charcoal A_s .
- the ^{surface} temperature of the charcoal T_s
- a constant of nature, σ , the Stefan-Boltzmann constant

$$\sigma = 5.67 \times 10^{-8} \frac{\text{W}}{\text{m}^2 \text{K}^4}$$

The units of the constant telegraph the equation

$$\dot{q}_r = \sigma A T_s^4$$

Place the lump of charcoal in a room at temp T_{sur} and not only can the charcoal emit heat but it can also absorb it.

Detailed balance demands that both appear in the same way:

$$\dot{q}_r = \sigma A (T_s^4 - T_{\text{sur}}^4)$$

finally if you change the surface of the charcoal, paint it,

coat it w/ ice, wrap it in tin foil you will change the efficiency with which the heat can leave. the Emissivity

ϵ covers this \dot{q}

$$0 \leq \epsilon \leq 1$$

$$\dot{q}_r = \epsilon \sigma A (T_s^4 - T_{\text{sur}}^4)$$

lowest values of ϵ happen for shiny metal (tin foil $\epsilon \sim 0.05$)

and increase from
then clean, or polished
metal

$\epsilon \sim 0.15$

rusty metal

$\epsilon \sim 0.5 \pm 0.15$

glass ~ 0.9

water, skin
 ~ 0.93

special paint
(high ϵ)

$\epsilon \sim 0.96$